

Bioengineering 280A
Principles of Biomedical Imaging

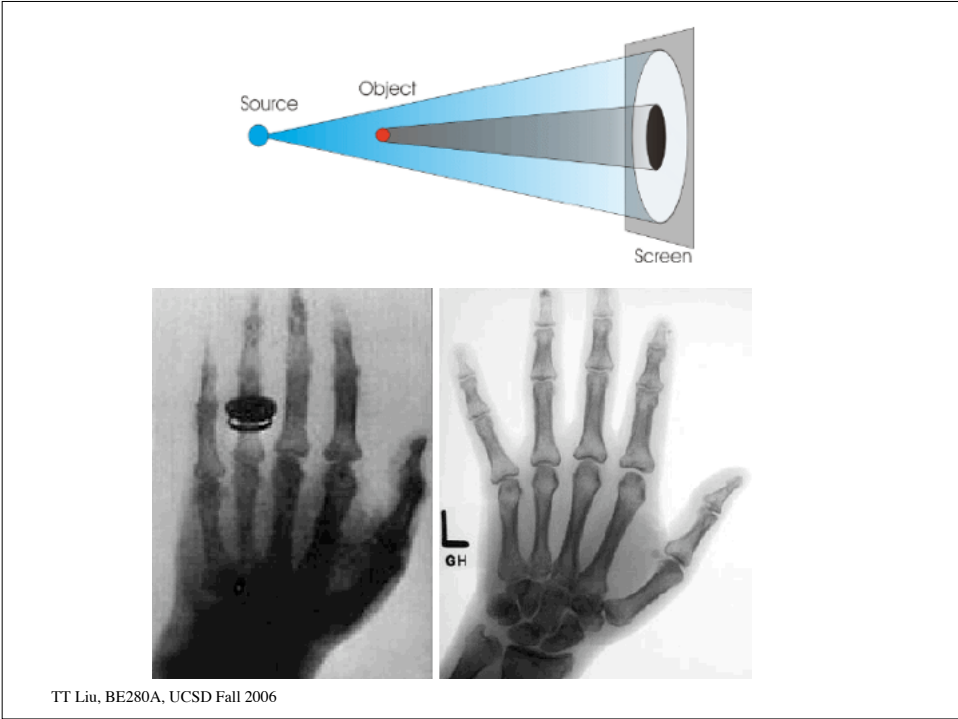
Fall Quarter 2006
X-Rays Lecture 2

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Topics

- Review topics from last lecture
- Attenuation
- Contrast
- Noise
- Image Equation

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X-Ray Production

Collisional transfers

Radiative transfers

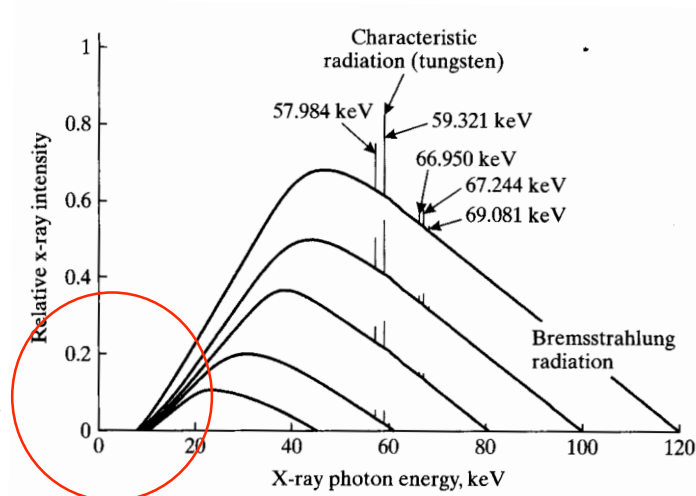
The diagram shows three stages of X-ray production:

- (a) **Collisional transfers:** An electron (represented by a circle with a minus sign) moves from left to right, passing through several atoms (represented by circles with a plus sign in the center). A 'Delta ray' is shown as a secondary electron ejected from one of the atoms.
- (b) **Radiative transfers:** An electron passes through an atom, causing an inner-shell electron to be ejected. This transition results in the emission of a 'Characteristic x-ray'.
- (c) **Radiative transfers:** An electron is decelerated as it passes near the 'Nucleus' of an atom, resulting in the emission of a 'Bremsstrahlung x-ray'.

Prince and Links 2005

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X-Ray Spectrum



Lower energy photons are absorbed by anode, tube, and other filters

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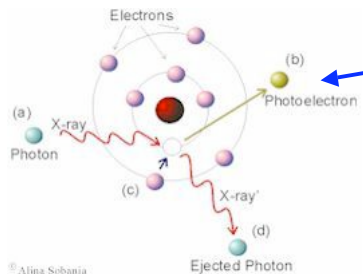
Prince and Links 2005

Interaction with Matter

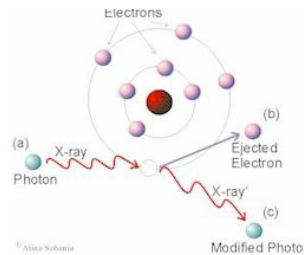
Typical energy range for diagnostic x-rays is below 200 keV.

The two most important types of interaction are photoelectric absorption and Compton scattering.

Photoelectric effect dominates at low x-ray energies and high atomic numbers. $E_{e^-} = h\nu - E_B$



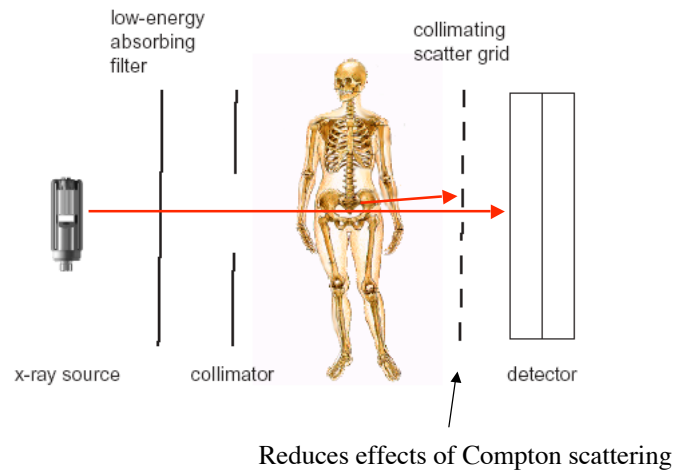
Compton scattering dominates at high x-ray energies and low atomic numbers, not much contrast



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<http://www.eee.ntu.ac.uk/research/vision/asobania>

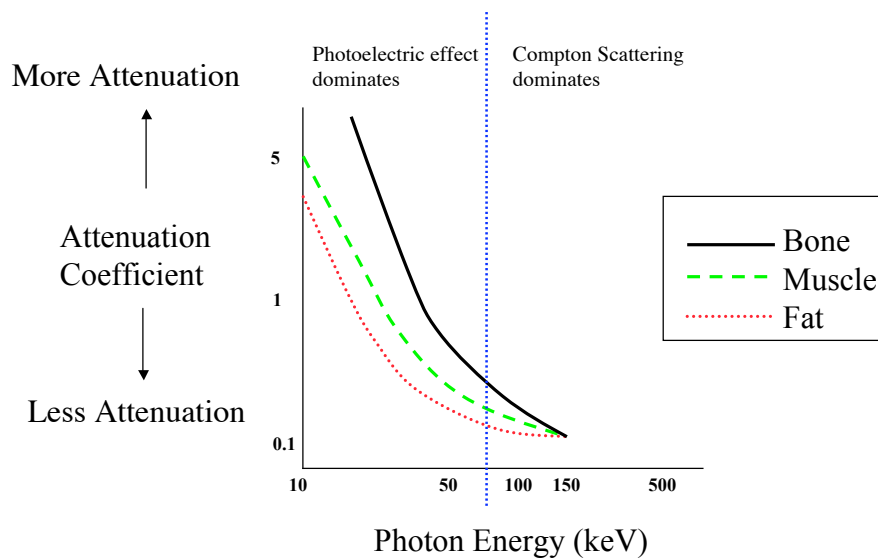
X-Ray Imaging Chain



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Suetens 2002

Attenuation



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Adapted from www.cis.rit.edu/class/simg215/xrays.ppt

Intensity

$$I = E\phi$$

Energy Photon flux rate

$$\phi = \frac{N}{A\Delta t}$$

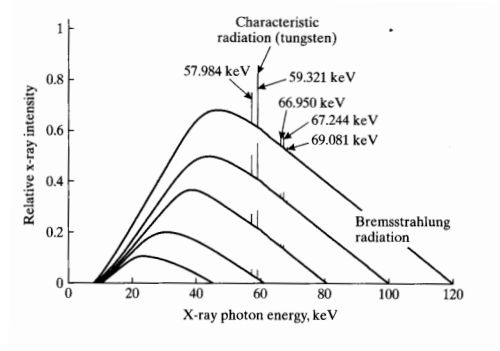
Number of photons
Unit Area Unit Time

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Intensity

$$\phi = \int_0^{\infty} S(E')dE'$$

X-ray spectrum



$$I = \int_0^{\infty} S(E')E'dE'$$

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Attenuation

$n = \mu N \Delta x$ photons lost per unit length

$\mu = \frac{n/N}{\Delta x}$ fraction of photons lost per unit length

$$\Delta N = -n \longrightarrow \frac{dN}{dx} = -\mu N \longrightarrow N(x) = N_0 e^{-\mu x}$$

For mono-energetic case, intensity is

$$I(\Delta x) = I_0 e^{-\mu \Delta x}$$

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Attenuation

Inhomogeneous Slab

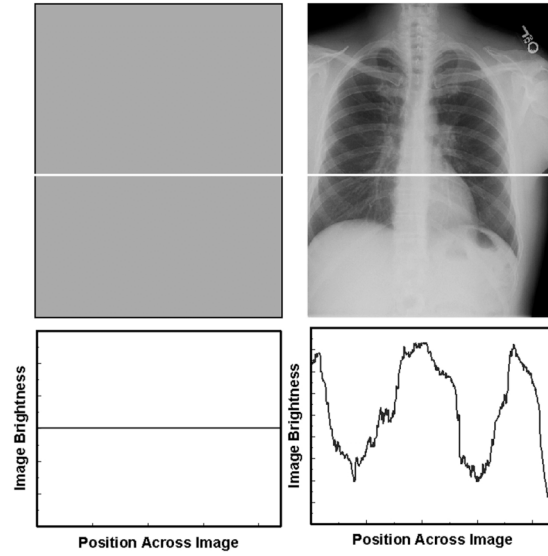
$$\frac{dN}{dx} = -\mu(x)N \longrightarrow N(x) = N_0 \exp\left(-\int_0^x \mu(x') dx'\right)$$
$$I(x) = I_0 \exp\left(-\int_0^x \mu(x') dx'\right)$$

Attenuation depends on energy, so also need to integrate over energies

$$I(x) = \int_0^\infty S_0(E') E' \exp\left(-\int_0^x \mu(x'; E') dx'\right) dE'$$

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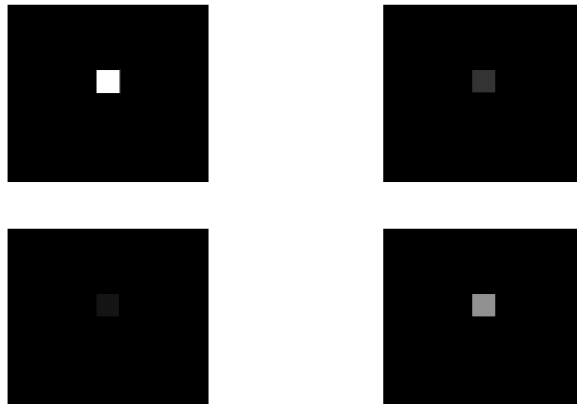
Contrast



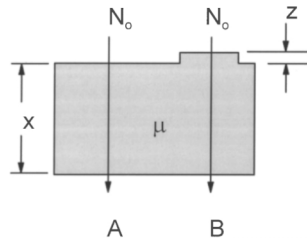
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Bushberg et al 2001

Contrast



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(A) X-ray Imaging

Background intensity

$$A = N_0 \exp(-\mu x)$$

Bushberg et al 2001

Object intensity

$$B = N_0 \exp(-\mu(x + z))$$

Subject/Local Contrast

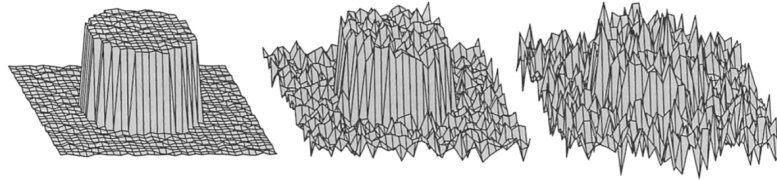
$$C_s = \frac{B - A}{A}$$

$$= \frac{N_0 \exp(-\mu(x + z)) - N_0 \exp(-\mu x)}{N_0 \exp(-\mu x)}$$

$$= \exp(-\mu z) - 1$$

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Noise and Image Quality



Low Noise

Medium Noise

High Noise

Bushberg et al 2001

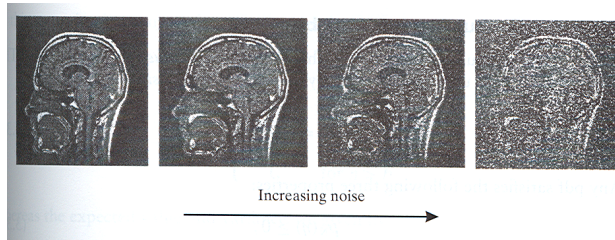


Figure 3.10

The effect of noise on image quality: image quality decreases rapidly with increasing noise contamination. Prince and Links 2005

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What is Noise?

Fluctuations in either the imaging system or the object being imaged.

Quantization Noise: Due to conversion from analog waveform to digital number.

Quantum Noise: Random fluctuation in the number of photons emitted and recorded.

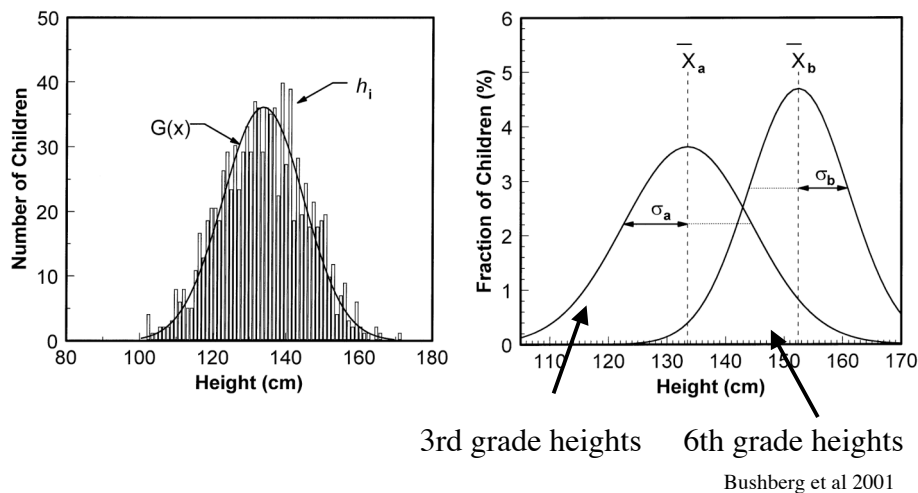
Thermal Noise: Random fluctuations present in all electronic systems. Also, sample noise in MRI

Other types: flicker, burst, avalanche - observed in semiconductor devices.

Structured Noise: physiological sources, interference

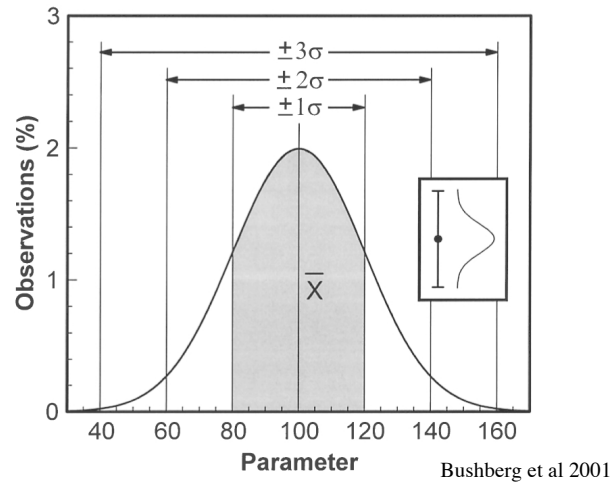
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Histograms and Distributions



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Gaussian Distribution



1, 2, and 3 standard deviation intervals correspond to 68%, 95%, and 99% of the observations

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Poisson Process

Events occur at random instants of time at an average rate of λ events per second.

Examples: arrival of customers to an ATM, emission of photons from an x-ray source, lightning strikes in a thunderstorm.

Assumptions:

- 1) Probability of more than 1 event in a small time interval is small.
- 2) Probability of event occurring in a given small time interval is independent of another event occurring in other small time intervals.

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Poisson Process

$$P[N(t) = k] = \frac{(\lambda t)^k}{k!} \exp(-\lambda t)$$

λ = Average rate of events per second

λt = Average number of events at time t

λt = Variance in number of events

Probability of interarrival times

$$P[T > t] = e^{-\lambda t}$$

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Example

A service center receives an average of 15 inquiries per minute. Find the probability that 3 inquiries arrive in the first 10 seconds.

$$\lambda = 15/60 = 0.25$$

$$\lambda t = 0.25(10) = 2.5$$

$$P[N(t = 10) = 3] = \frac{(2.5)^3}{3!} \exp(-2.5) = .2138$$

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Quantum Noise

Fluctuation in the number of photons emitted by the x-ray source and recorded by the detector.

$$P_k = \frac{N_0^k \exp(-N_0)}{k!}$$

P_k : Probability of emitting k photons in a given time interval.

N_0 : Average number of photons emitted in that time interval = λt

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Transmitted Photons

$$Q_k = \frac{(tN_0)^k \exp(-tN_0)}{k!}$$

Q_k : Probability of k photons making it through object

N_0 : Average number of photons emitted in that time interval = λt

$t = \exp(-\int \mu dz)$ = fraction of photons transmitted

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Mean and Variance

For a Poisson process, the mean = variance, i.e. $\bar{X} = \sigma^2$

Therefore, the standard deviation is given by $\sigma = \sqrt{\bar{X}}$

For X-ray systems, if the mean number of counts is N, then the standard deviation in the number of counts is $\sigma = \sqrt{N}$.

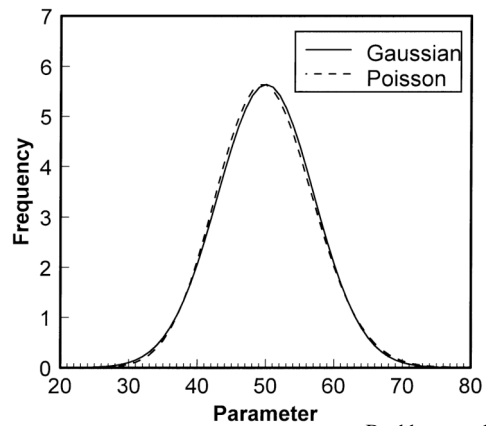
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TABLE 10-1. EXAMPLES OF NOISE VERSUS PHOTONS

Photons/Pixel (N)	Noise (σ) ($\sigma = \sqrt{N}$)	Relative Noise (σ/N) (%)	SNR (N/σ)
10	3.2	32	3.2
100	10	10	10
1,000	31.6	3.2	32
10,000	100	1.0	100
100,000	316.2	0.3	316

SNR, signal-to-noise ratio.

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Bushberg et al 2001

Poisson Distribution describes x - ray counting statistics.

Gaussian distribution is good approximation to Poisson when $\sigma = \sqrt{\bar{X}}$

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- G79.** A series of measurements has a mean of 100 counts. A range of $\pm\sigma$ is ____ .
- A. 95–105
 - B. 90–110
 - C. 68–137
 - D. 50–150
 - E. 33–167

- G80.** To achieve a standard deviation of 2%, ____ counts must be collected.
- A. 400
 - B. 1,414
 - C. 2,500
 - D. 10,000
 - E. 40,000

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- G73.** A radioactive sample is counted many times, and the mean is 2500 counts. 96% of the readings will lie between _____ and _____ counts.
- A. 2300 2500
 - B. 2400 2500
 - C. 2400 2600
 - D. 2450 2550
 - E. 2500 2700

- D70.** How many counts must be collected in an instrument with zero background to obtain an error limit of 1% with a confidence interval of 95%?
- A. 1000
 - B. 3162
 - C. 10,000
 - D. 40,000
 - E. 100,000

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Contrast and SNR for X-Rays

$$\text{Contrast} = C = \frac{I_t - I_b}{I_b}$$

$$\text{SNR} = \frac{I_t - I_b}{\sigma_b}$$

$$I_b = \frac{N_b \cdot E}{A\Delta t} \rightarrow \text{var}(I_b) = \text{var}(N_b) \left(\frac{E}{A\Delta t} \right)^2 = N_b \left(\frac{E}{A\Delta t} \right)^2$$

$$\sigma_b = \text{std}(I_b) = \sqrt{N_b} \left(\frac{E}{A\Delta t} \right)$$

$$\text{SNR} = \frac{CI_b}{\sigma_b} = C\sqrt{N_b} = C\sqrt{\Phi A R t \eta}$$

Photons/Roentgen/cm²
Area
Exposure in Roentgens
Detector efficiency
Fraction transmitted

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Example

$$\Phi = 637 \times 10^6 \text{ photons R}^{-1}\text{cm}^{-2}$$

$$R = 50 \text{ mR}$$

$$t = 0.05$$

$$\eta = 0.25$$

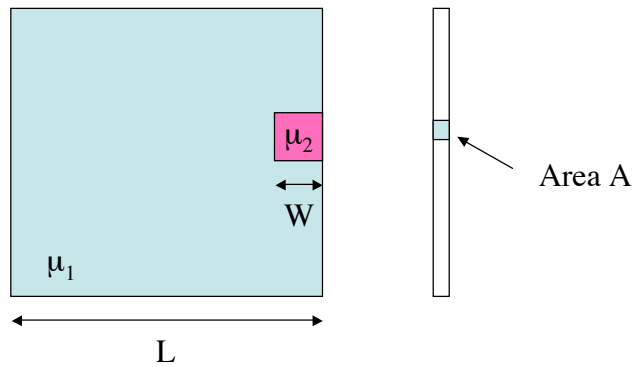
$$A = 1\text{mm}^2$$

$$C = 0.1 \text{ (10\% contrast)}$$

$$\text{SNR} = 0.1\sqrt{6.37 \times 10^8 \cdot .05 \cdot .25 \cdot .01} = 6.3$$

$$20\log_{10}(6.3) = 16 \text{ dB}$$

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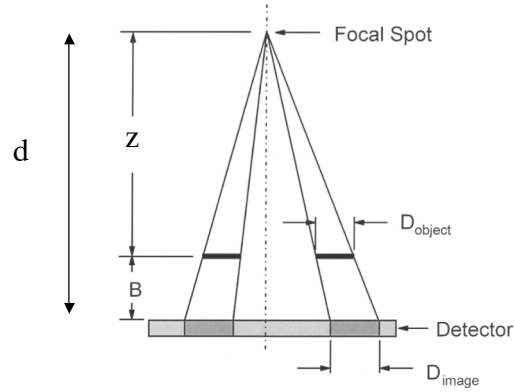


$$C = \frac{I_t - I_b}{I_b} = \frac{N_0 \left(\exp(-(\mu_1(L - W) + \mu_2 W)) - \exp(-\mu_1 L) \right)}{N_0 \exp(-\mu_1 L)}$$

$$\text{SNR} = C \sqrt{N_0 A \exp(-\mu_1 L)}$$

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Magnification of Object



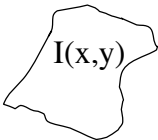
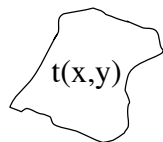
$$M(z) = \frac{d}{z}$$

$$= \frac{\text{Source to Image Distance (SID)}}{\text{Source to Object Distance (SOD)}}$$

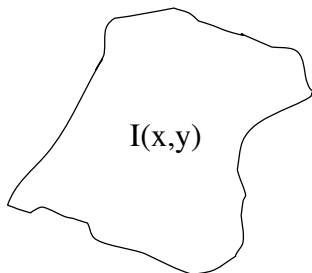
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Bushberg et al 2001

Magnification of Object



$$M = 1: I(x,y) = t(x,y)$$



$$M = 2: I(x,y) = t(x/2, y/2)$$

$$\text{In general, } I(x,y) = t(x/M(z), y/M(z))$$

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Bushberg et al 2001

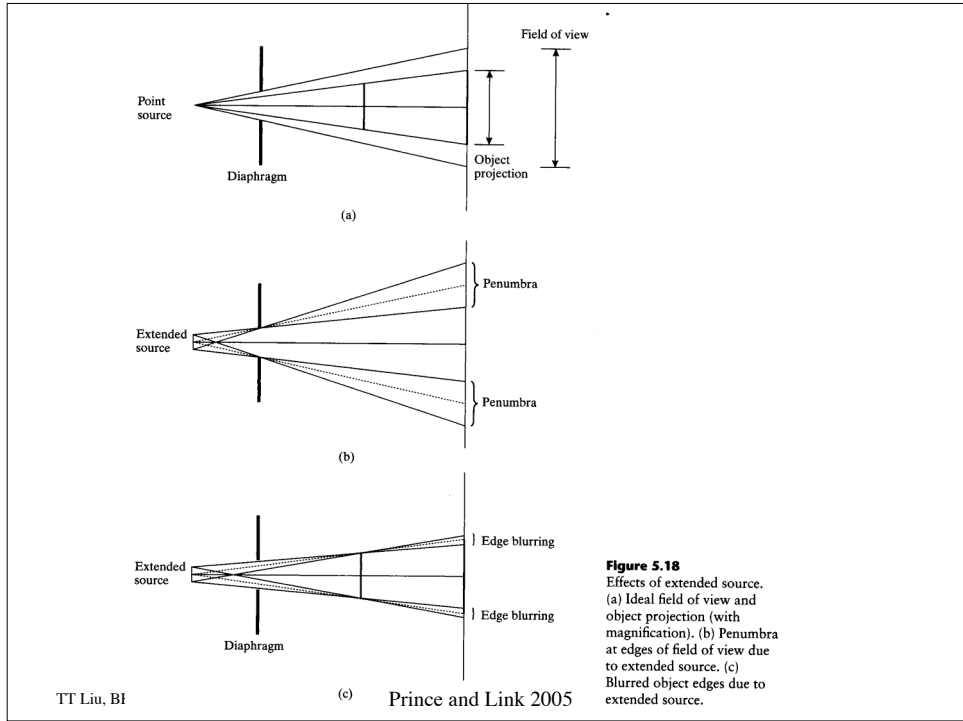


Figure 5.18
Effects of extended source. (a) Ideal field of view and object projection (with magnification). (b) Penumbra at edges of field of view due to extended source. (c) Blurred object edges due to extended source.

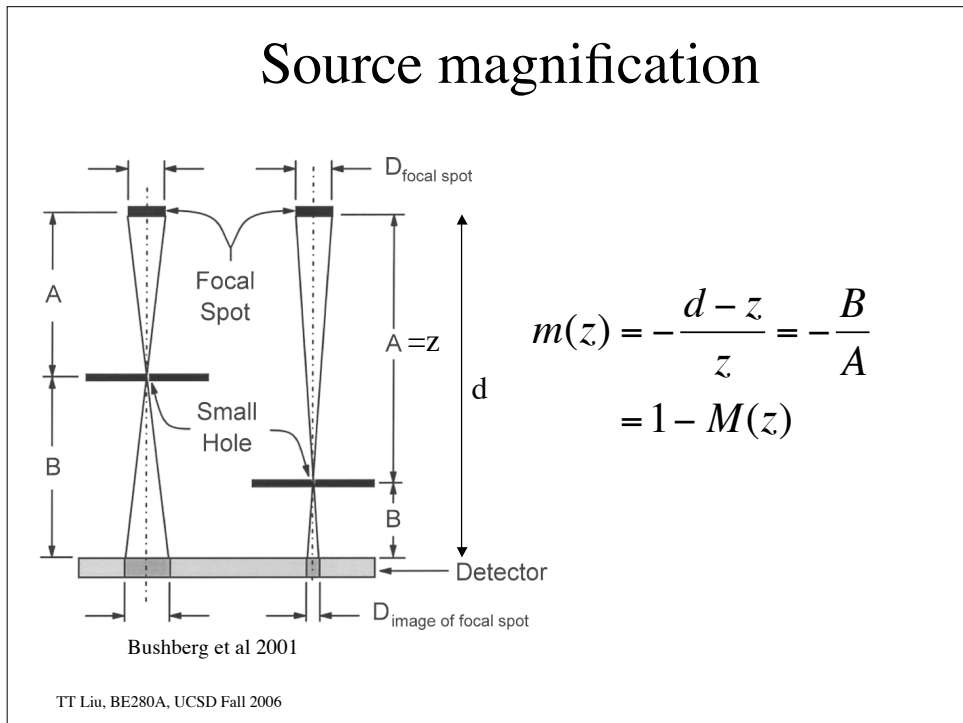


Image of a point object

$s(x,y)$



$$I_d(x,y) = \lim_{m \rightarrow 0} ks(x/m, y/m)$$

$$= \delta(x,y)$$

$s(x,y)$



$m=1$



$$I_d(x,y) = ks(x,y)$$

In general,

$$I_d(x,y) = ks(x/m, y/m)$$

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Image of arbitrary object

$s(x,y)$



$t(x,y)$



$$\lim_{m \rightarrow 0} I_d(x,y) = t(x,y)$$

$s(x,y)$



$t(x,y)$



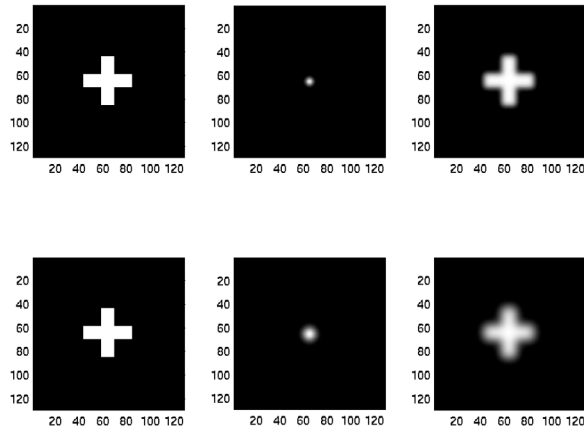
$m=1$

$$I_d(x,y) = ???$$

$$I_d(x,y) = ks(x/m, y/m) ** t(x/M, y/M)$$

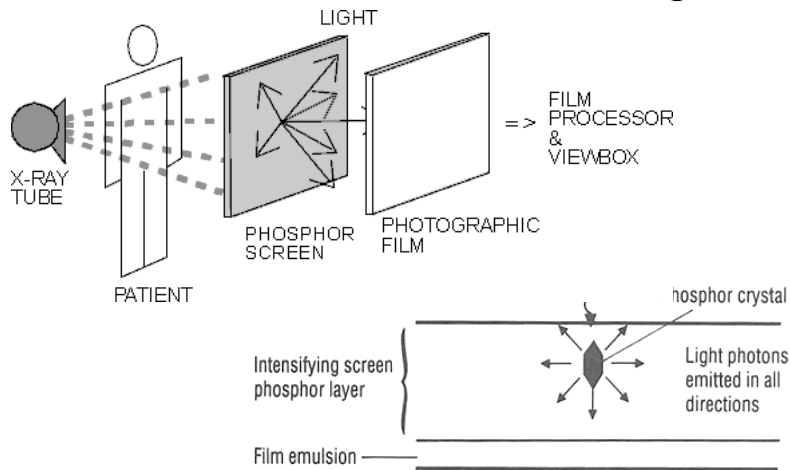
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Convolution



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Film-screen blurring



$$I_d(x, y) = ks(x/m, y/m) ** t(x/M, y/M) ** h(x, y)$$

http://learntech.uwe.ac.uk/radiography/RScience/imaging_principles_d/diagram11.htm
<http://www.sunnybrook.utoronto.ca:8080/~selenium/xray.html#Film>

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